

TOWARDS AN INTERFEROMETRIC AUTOFOCUS FOR IONOSPHERIC PHASE SIGNATURES IN BIOMASS

Felipe Betancourt-Payan*, Marc Rodriguez-Cassola, Pau Prats-Iraola, Maria J. Sanjuan-Ferrer,
Gerhard Krieger

Microwaves and Radar Institute, German Aerospace Center (DLR)



Table of Contents

Biomass and the Ionospheric Errors

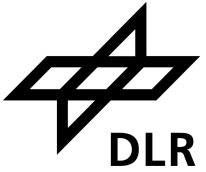
Calibration Errors
Characterization

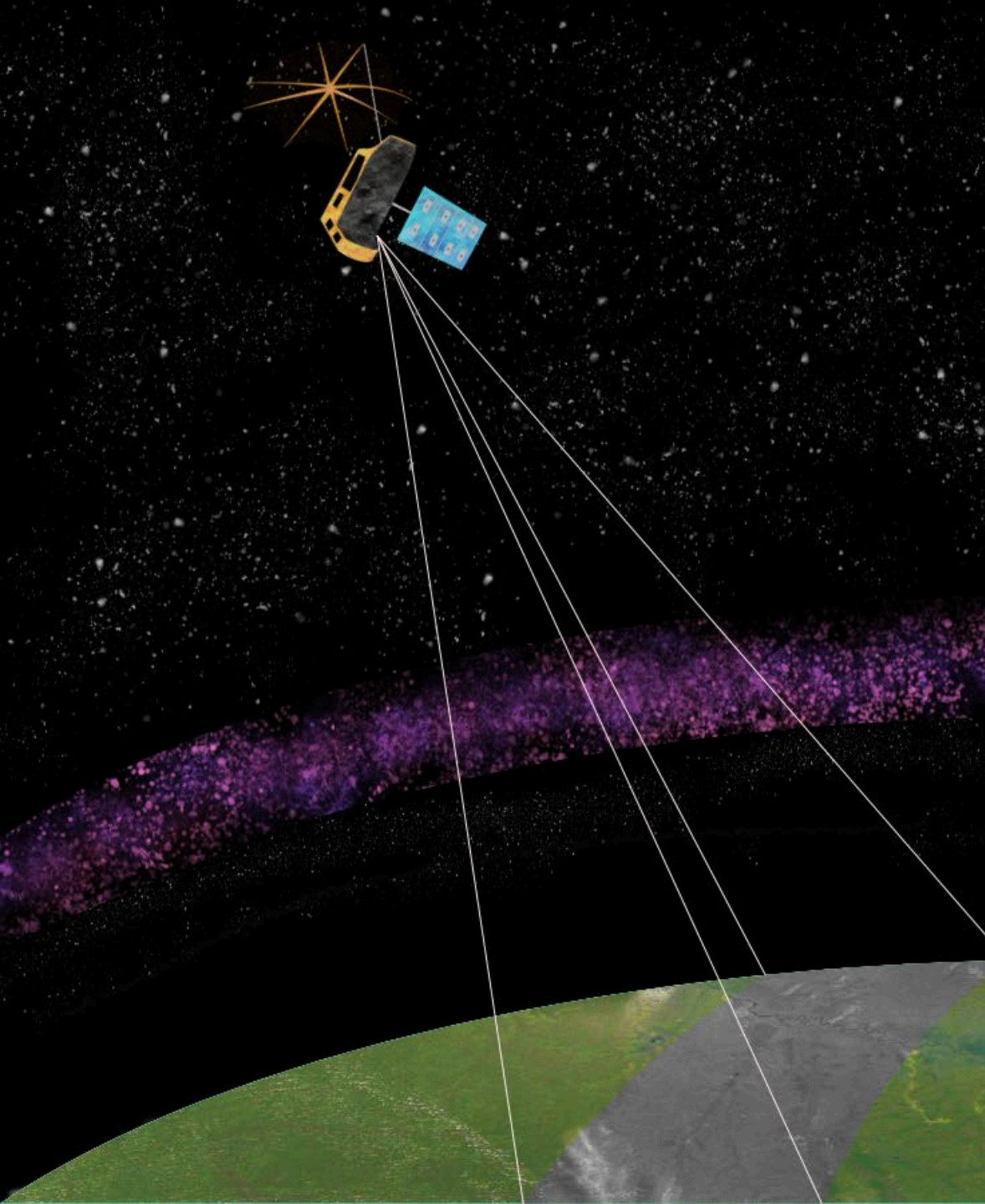
Faraday Rotation

Autofocus

Interferometric
Data Combination

Conclusions





Biomass

- Launch: 2025
- P-band (438 MHz)
- Polarimetry: Quad-pol
- Diameter of antenna: 12 m

Scientific Objectives

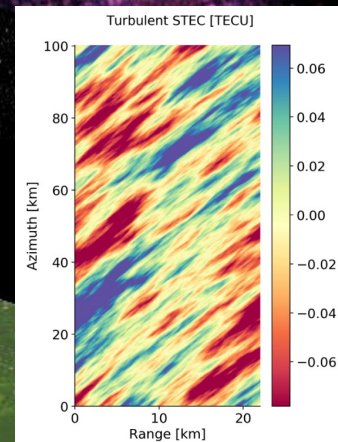
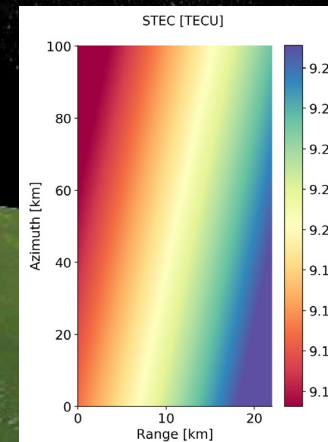
- Measure global biomass and contribution of change in carbon fluxes
- Subsurface geology
- Terrain topography under dense vegetation
- Glacier and icesheet velocities

Ionospheric effects on products

- Delay of wavefront: geolocation errors
- Phase screens
- Phase errors: azimuth shifts and defocusing
- Dispersion: broadening of IR
- Faraday rotation: polarization errors and channel imbalance
- Absorption

Background

Turbulent



Error Characterization in Ionospheric Calibration of SAR Images

Thin layer approximation

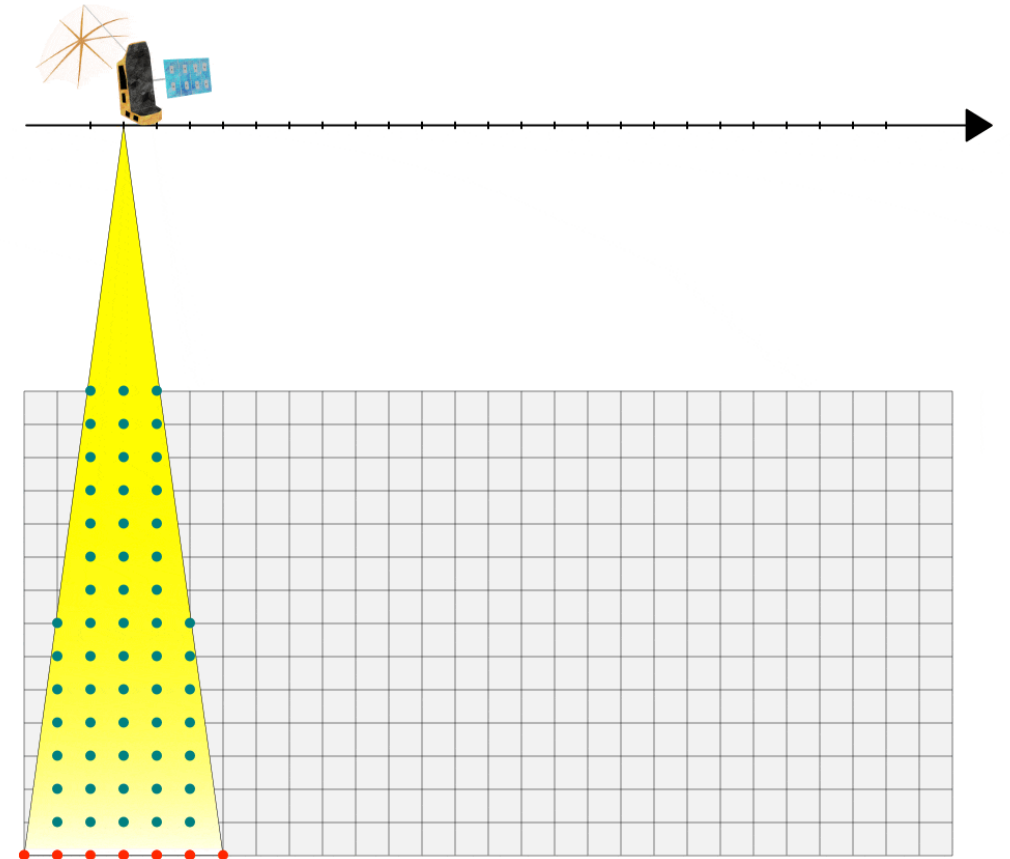
Background component can be removed

Irregularities are stochastic and stationary

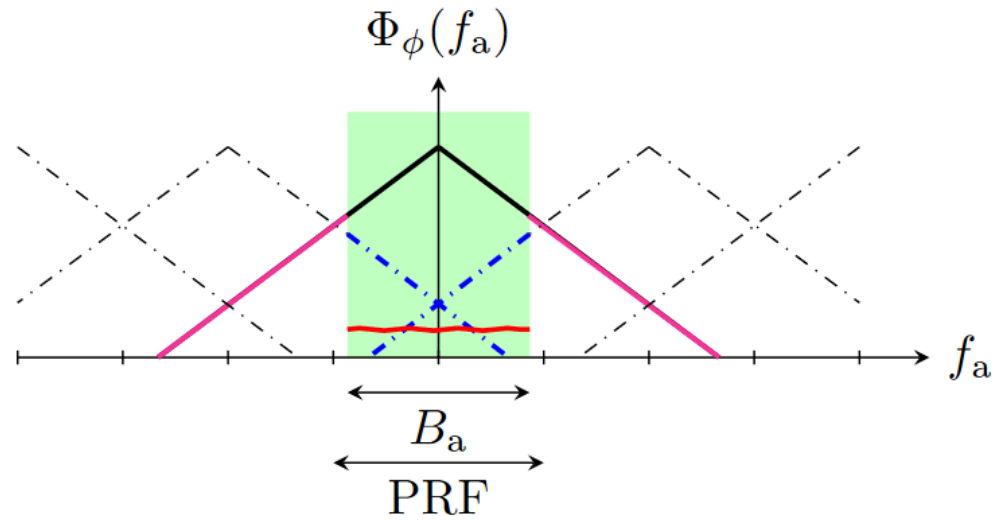
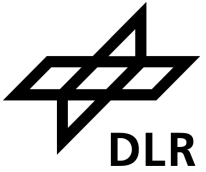
Rino power law

$$\Phi_{\phi}(\kappa_x, \kappa_y) = \frac{\lambda^2 \cdot r_e^2 \cdot \sec^2 \theta_{\text{inc}} \cdot ab \cdot \left(2 \cdot \pi / 1000\right)^{p+1} \cdot C_k L}{\left(\kappa_0^2 + A \cdot \kappa_x^2 + B \cdot \kappa_x \cdot \kappa_y + C \cdot \kappa_y^2\right)^{p+1/2}}$$

$$\Phi_{\Omega}(\kappa_x, \kappa_y) = \left(\frac{q_e \cdot \vec{B} \cdot \hat{k}}{m_e \cdot f \cdot 4 \cdot \pi}\right)^2 \cdot \Phi_{\phi}(\kappa_x, \kappa_y)$$

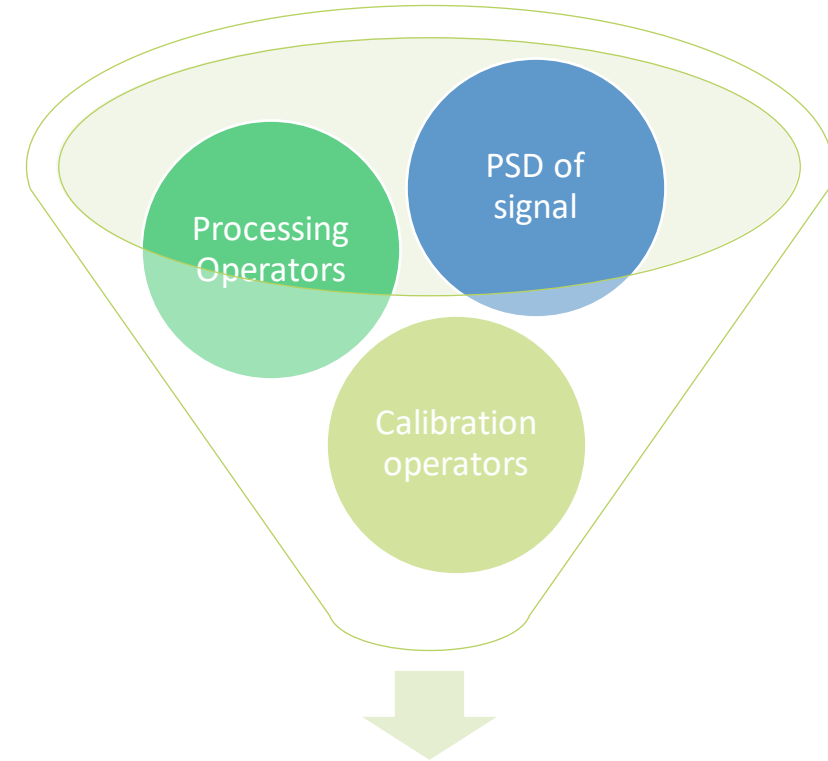


Calibration error sources



$$\Phi_{\Delta\phi}(f_a) = \begin{cases} \Phi_{\phi}(f_a), & f_a > B_a/2 \\ \Phi_N(f_a) + \sum_{i=1}^{+\infty} (\Phi_{\phi}(f_a + i \cdot \text{PRF}) + \Phi_{\phi}(f_a - i \cdot \text{PRF})), & f_a \leq B_a/2 \end{cases}$$

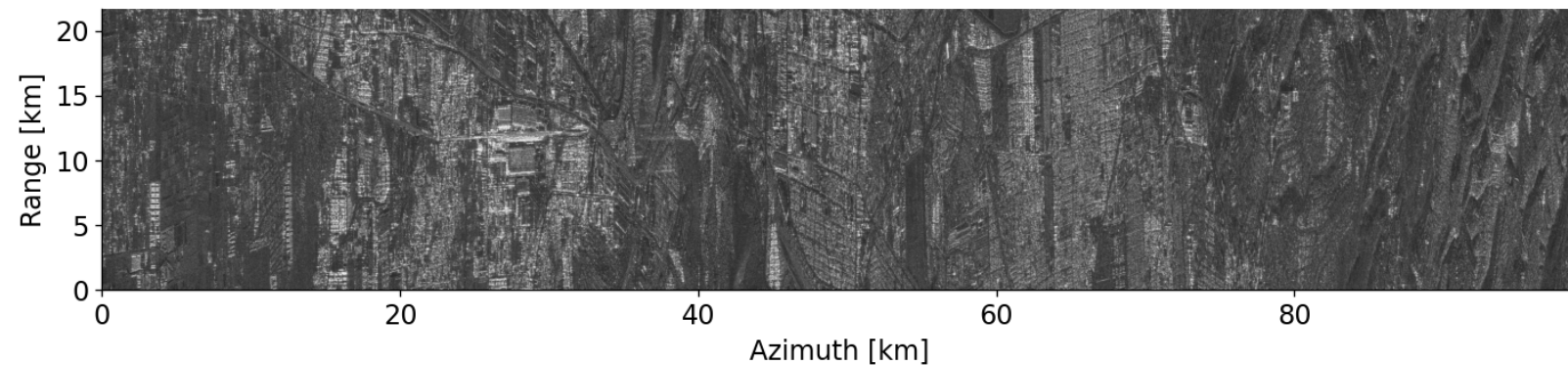
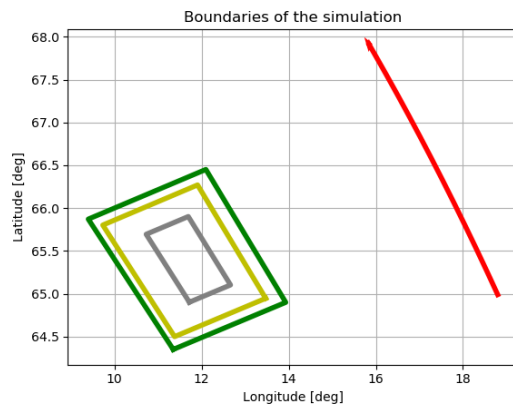
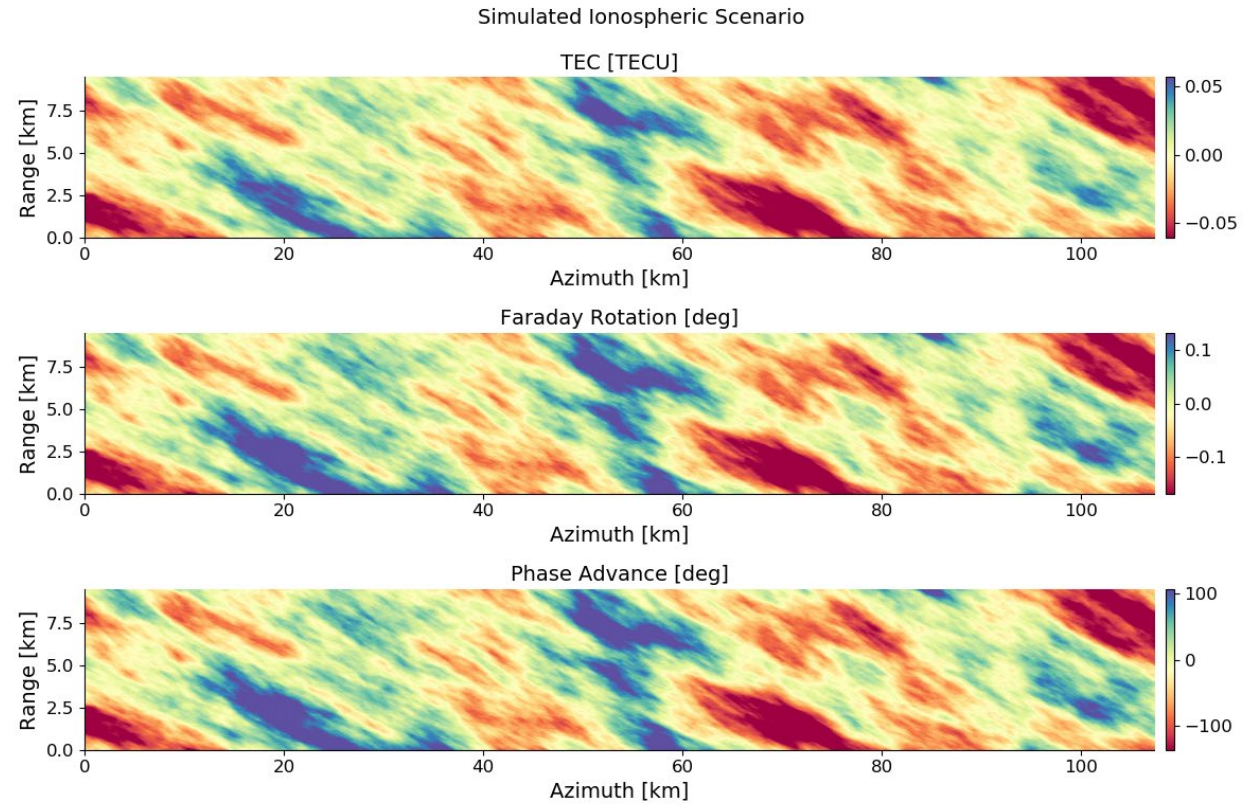
$$\sigma_{\Delta\phi}^2 = 2 \cdot \left(\underbrace{\int_0^{B_a/2} \Phi_N(f_a) df_a}_{\text{Noise}} + \underbrace{\sum_{i=1}^{+\infty} (\Phi_{\phi}(f_a + i \cdot \text{PRF}) + \Phi_{\phi}(f_a - i \cdot \text{PRF}))}_{\text{Aliasing}} \right) + \underbrace{\int_{B_a/2}^{\infty} \Phi_{\phi}(f_a) df_a}_{\text{Outside of band}}$$



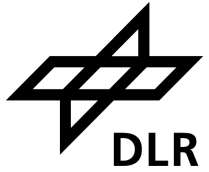
Expected solution and error

Data Simulation

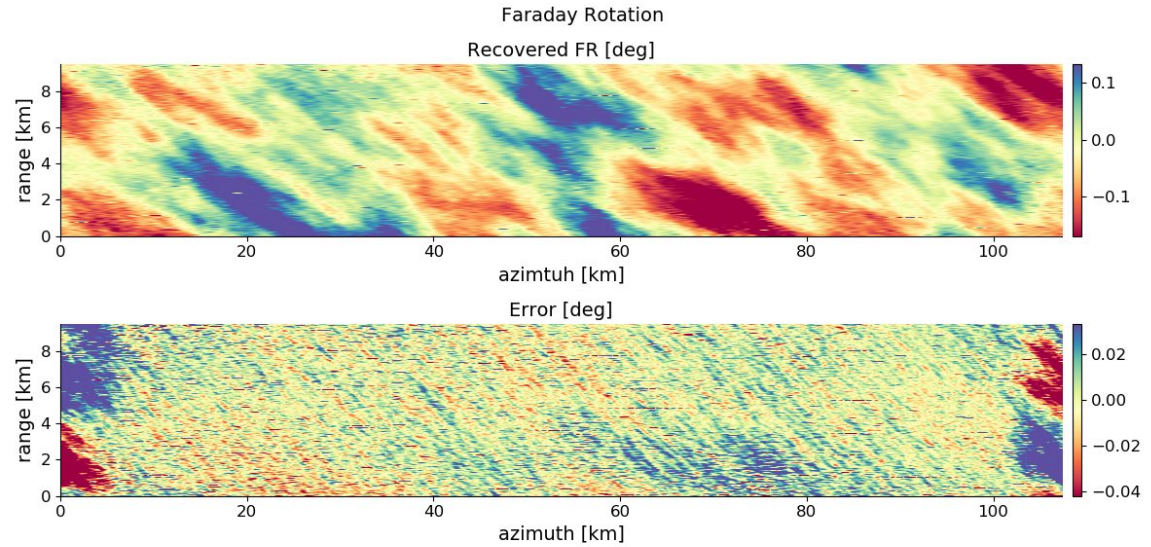
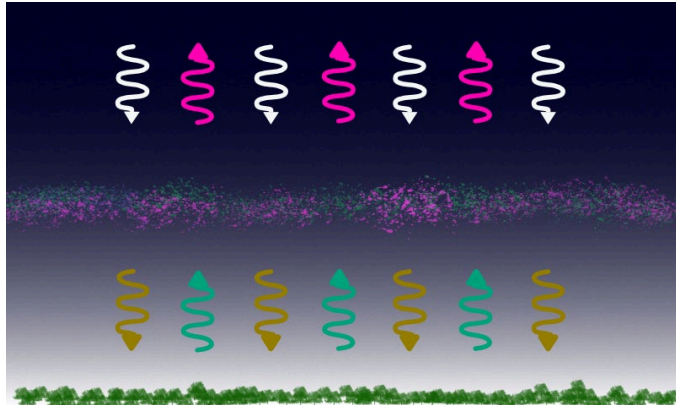
$C_k L$	10^{32}
Outer scale l_o	20 km
Anisotropy $a : b$	5:1
Spectral index p	2.65
Ionospheric height h_{iono}	351 km
Satellite altitude	660 km
Incidence angle	25 deg
Antenna diameter	12 m
Carrier frequency	435 MHz
Pulse Repetition Frequency (PRF)	1475.506 Hz
Azimuth bandwidth (B_a)	1229.588 Hz
Range sampling frequency (RSF)	7565217.4 Hz
Range bandwidth (B_r)	6877470.363 Hz



Bickel and Bates Based Estimation

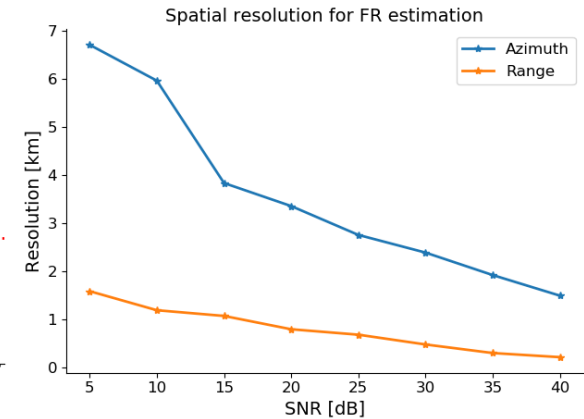
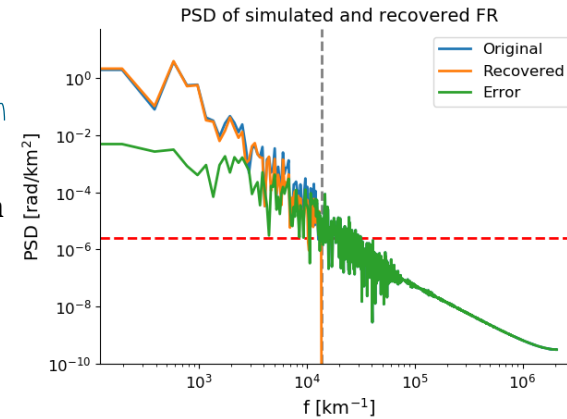


(Tx) H, V → (Rx) HH, HV, VH, VV

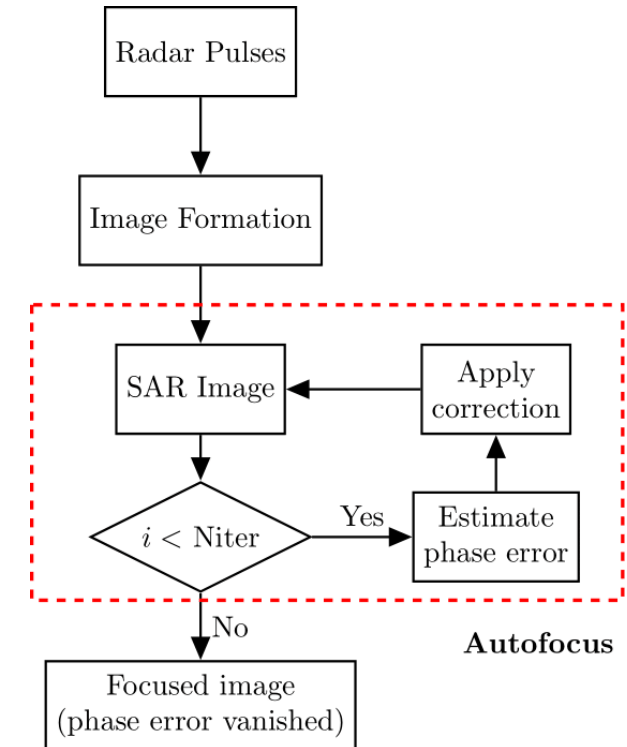
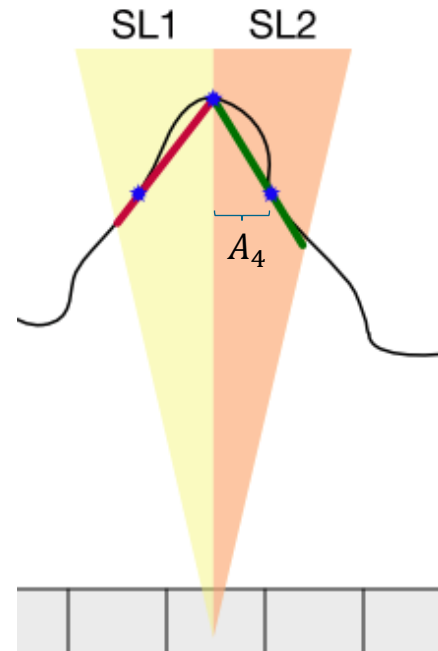
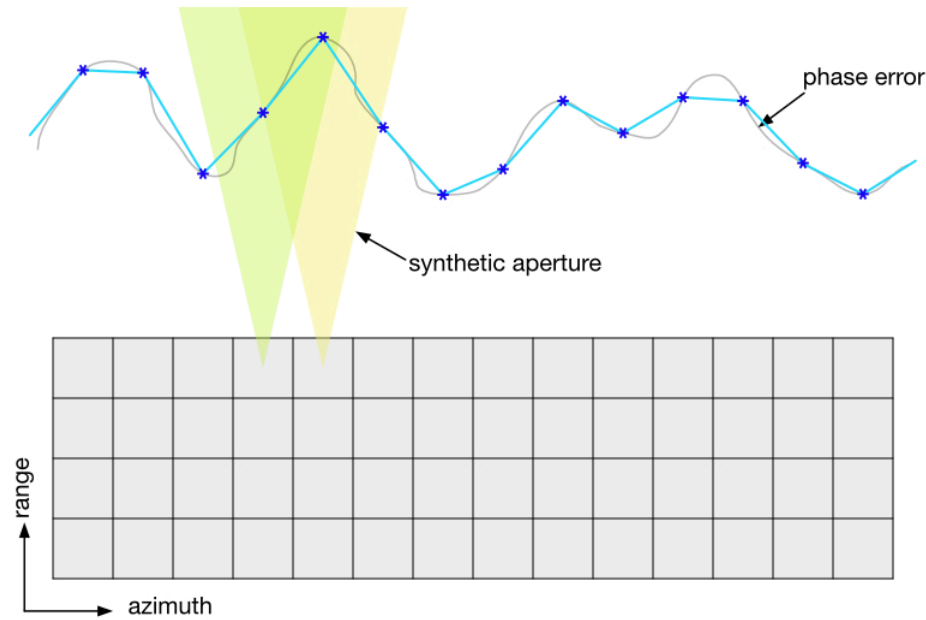


$$\sigma_{\Delta\Omega}^2 = \iint_{S_{cbw}} \underbrace{B^2 \cdot \Phi_N}_{\text{Filtered noise}} + \underbrace{\sum_{i=-\infty}^{+\infty} \sum_{j=-\infty}^{\infty} \Phi_{\Omega}[i,j]}_{\text{Aliasing}} df_r \cdot df_a + \iint_{S_{cbw}^c} \Phi_{\Omega} df_a \cdot df_a$$

Inside of band
Outside of band



Mapdrift Autofocus



$$\Phi_{\phi, AF} = L^2 \cdot \frac{1}{(j \cdot \omega_2)^4} \left(\sum_{i=-\infty}^{+\infty} \sum_{j=-\infty}^{+\infty} \left(B^2 \cdot \Phi_{\phi^*}[i, j] \right) \right)$$

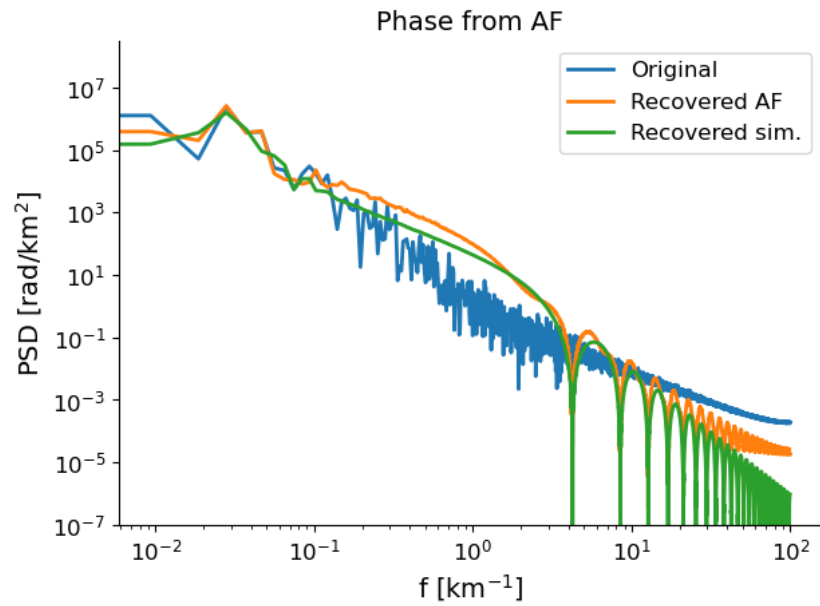
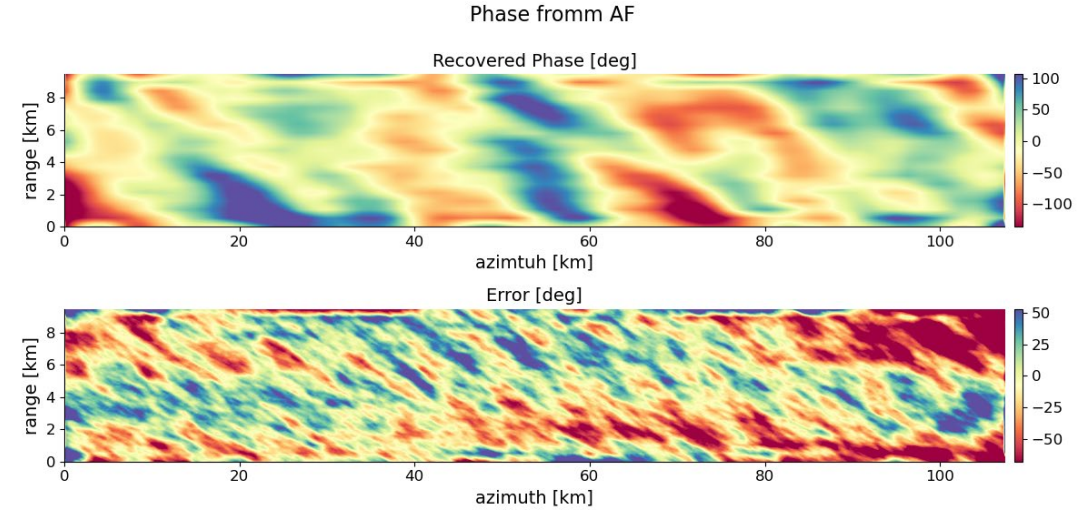
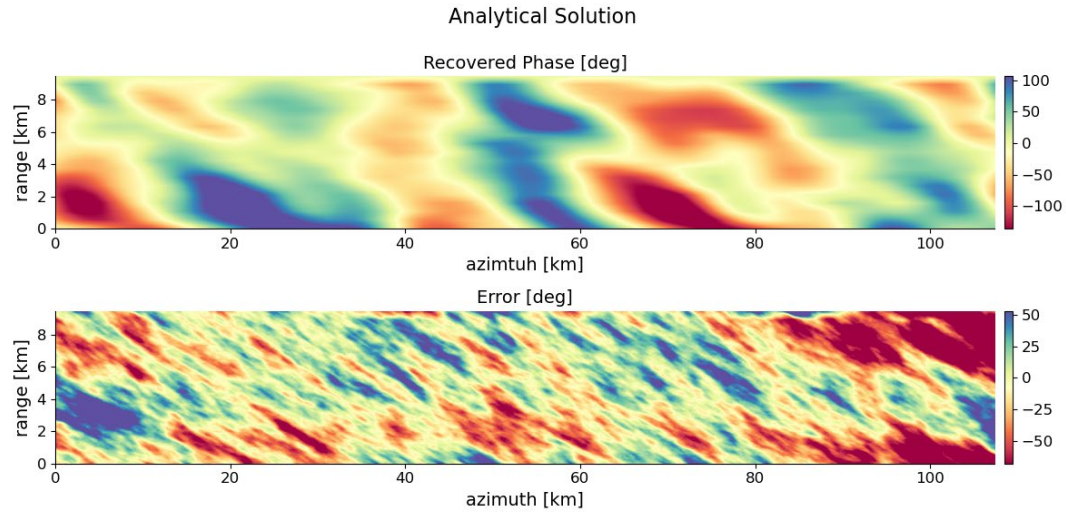
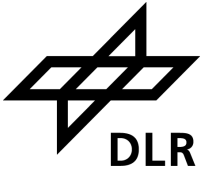
Interpolation
Integration
Sampling

Block averaging
Differentiation + Synthetic Aperture

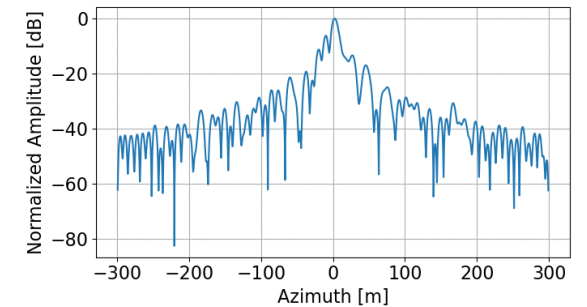
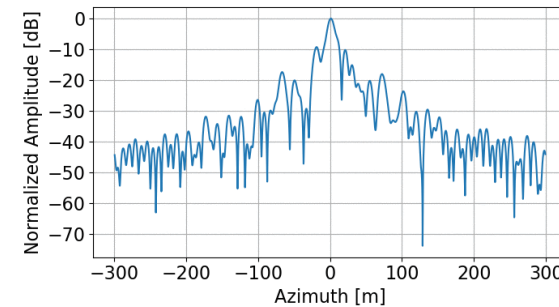
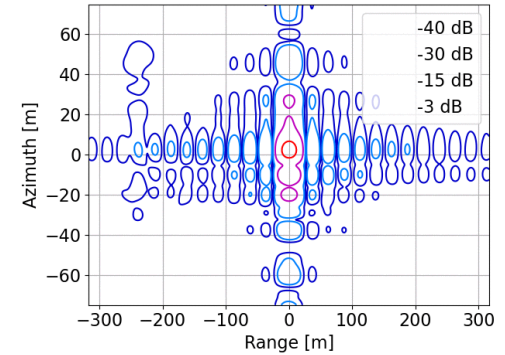
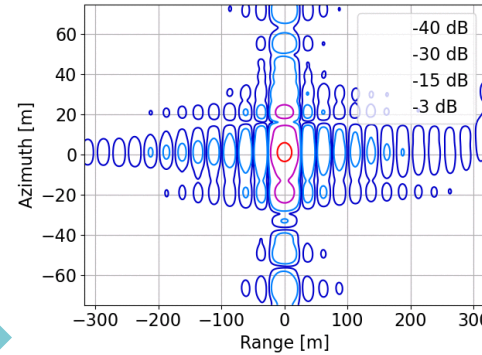
$$\ddot{\phi}^* = \exp(j2\pi f A_4) (\exp(j2\pi f A_4) \ddot{\phi} - \phi) - (\exp(j2\pi f A_4) \phi - \phi)$$

$$\Phi_{\ddot{\phi}^*} = \ddot{\phi}^* \cdot \overline{\ddot{\phi}^*}$$

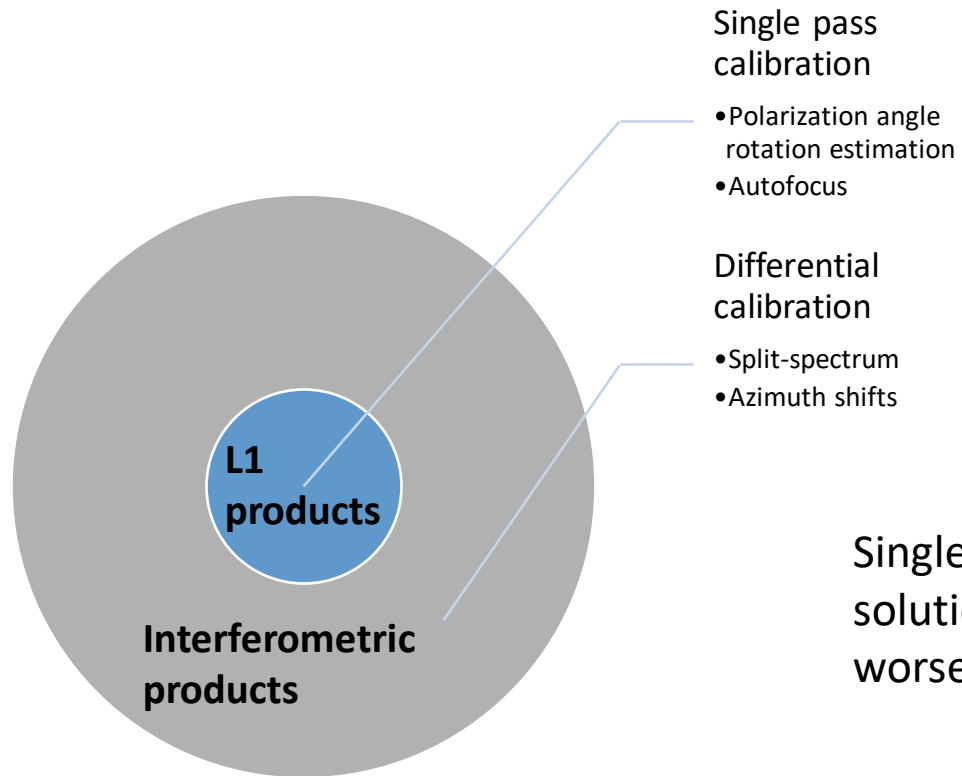
Mapdrift of Autofocus



IRF
with/without AF



Methods for Correcting Ionospheric Effects in SAR Products



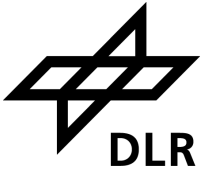
Interferometric calibration has better resolution but provides only differential ionospheric products

Single pass solutions have worse resolution ?

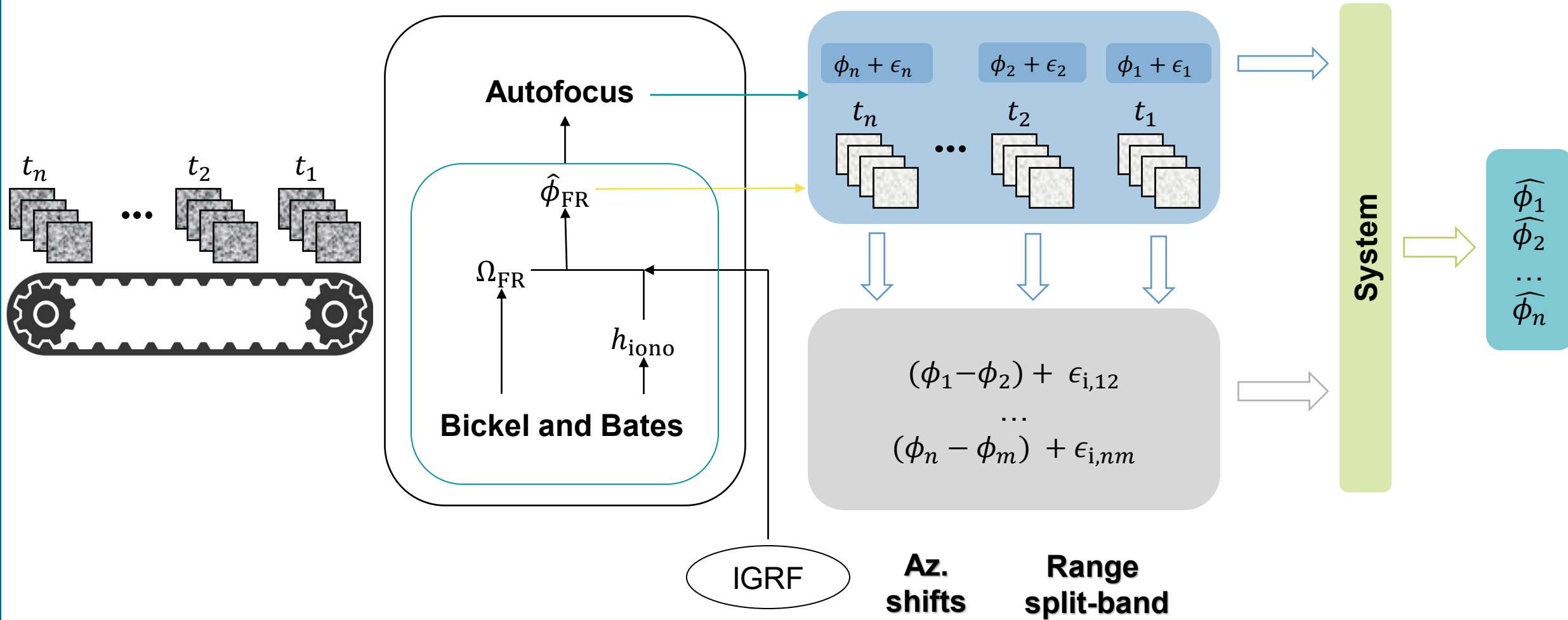


Can I profit from the interferometric information to search for better single pass solutions?

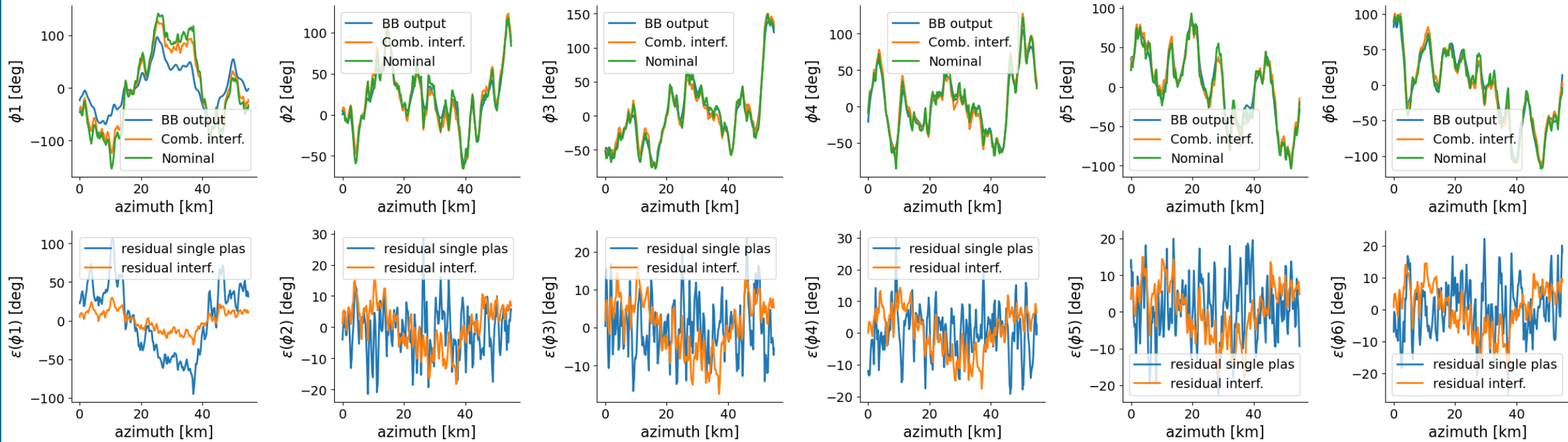
Data combination



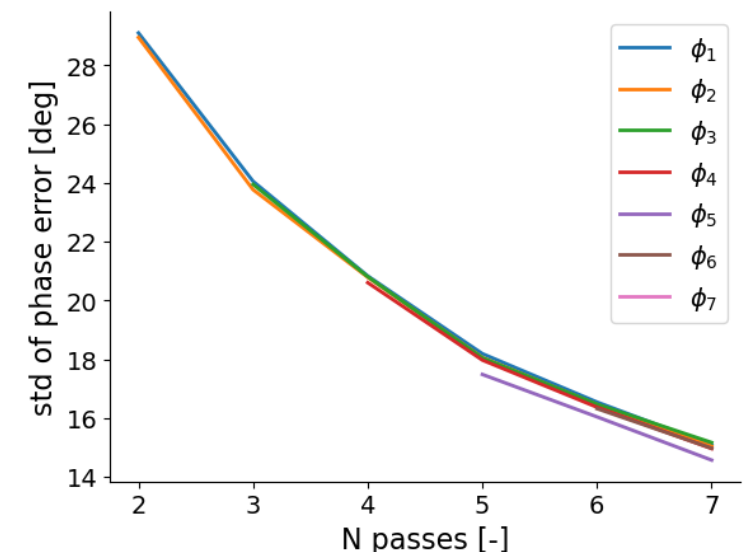
Ionospheric Calibration



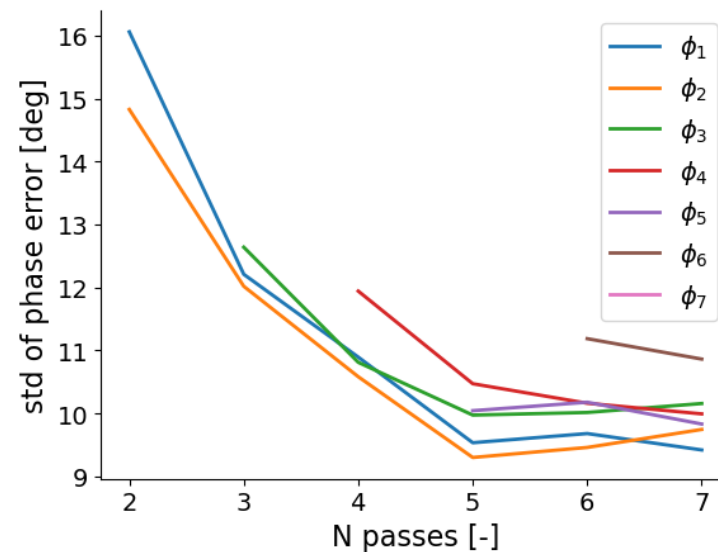
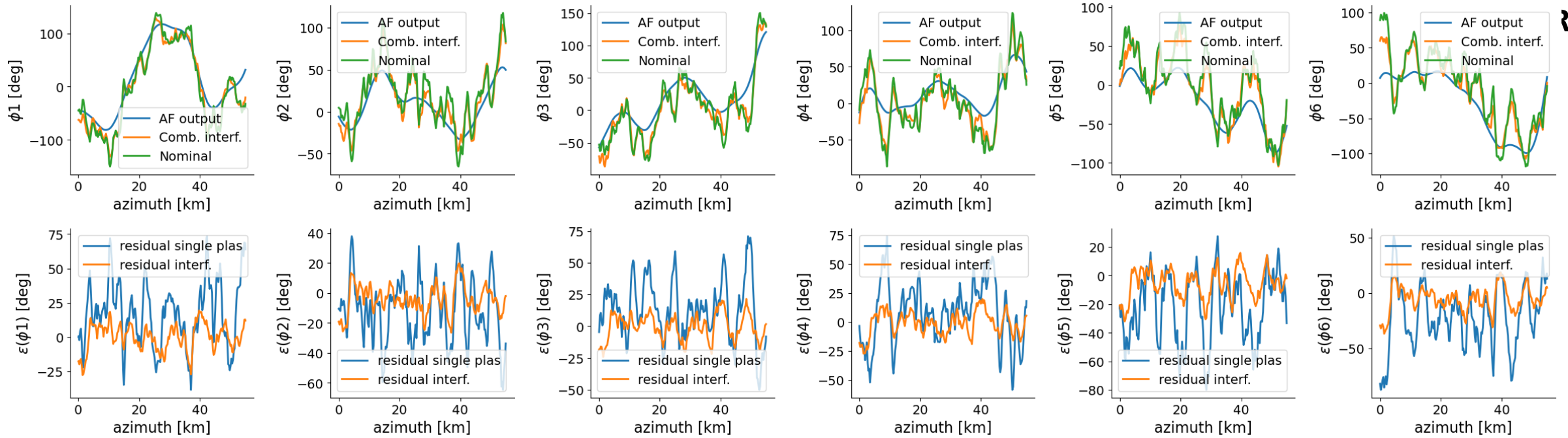
Combination of FR Outputs



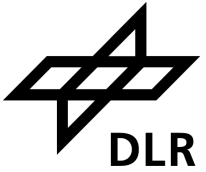
- Combining measurements and forcing the individual solutions to be consistent with the interferograms it is possible to gain resolution in the ionospheric maps
- The more information we have the better
- Remember: only suitable in high latitudes!



Combination of Autofocus Outputs



Conclusions



Spectral analysis can be used for the characterization of the errors in the calibration algorithms (FR based and Autofocus)

Residual phase errors after single pass calibrations introduce large errors in the interferograms


Information obtained by the interferometric processing can be used to gain resolution and reduce uncertainties

The approach passes by forcing the single pass solutions to be consistent with the interferometric solutions.

Uncertainty decreases with number of passes



felipe.betancourtpayan@dlr.de

 @fe_betancourtp

 felipe.betancourtp